



= 3.0 GeV, corresponding to baryon chemical potentials from 420 MeV to about 720 MeV. The directed flow ( $v_1$ ) for protons and  $\Lambda$  hyperons at 7.7 GeV and above [3, 4] suggests a qualitative resemblance to the “softest point” of the equation of state predicted by hydrodynamic models.

Net-proton higher moment measurements from energies below 7.7 GeV could help determine the type of phase transition or reveal evidence for critical fluctuations [5]. However, RHIC cannot operate in collider mode below 7 GeV, due to steeply decreasing luminosity. By inserting a target into the beam pipe and circulating one beam in RHIC, we can study FXT collisions below  $\sqrt{s_{NN}} = 7$  GeV. Figure 1 shows an example of a schematic phase diagram, and illustrates the possible region probed by FXT measurements.

During a brief test in 2015, STAR collected approximately 1.3 million FXT events with centrality 0-30%.

## 2. Results

### 2.1. STAR FXT Particle Yield Results

Figure 2 presents proton spectra for 0-5% centrality in Au+Au collisions at  $\sqrt{s_{NN}} = 4.5$  GeV as a function of transverse mass  $m_T - m_p$  in several rapidity bins, each with a width  $\Delta y = 0.1$ . These spectra were fitted using the Blast-Wave model, after correcting for detector efficiency, acceptance, energy loss and hadronic background. Overall, fits describe the data across the STAR FXT acceptance range.

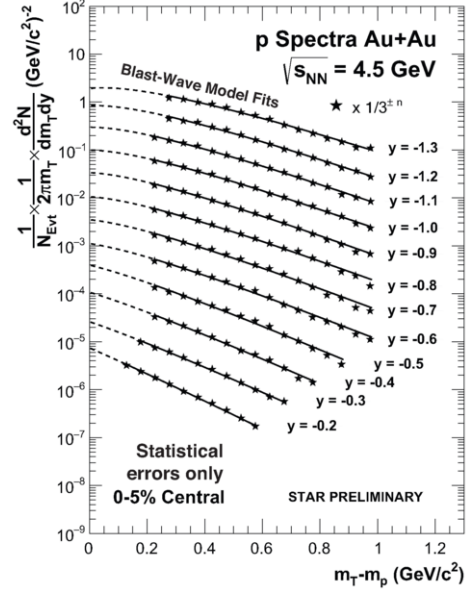


Fig. 2: Proton transverse mass spectra in 4.5 GeV Au+Au collisions.

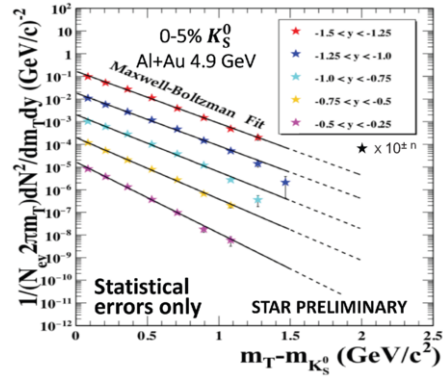
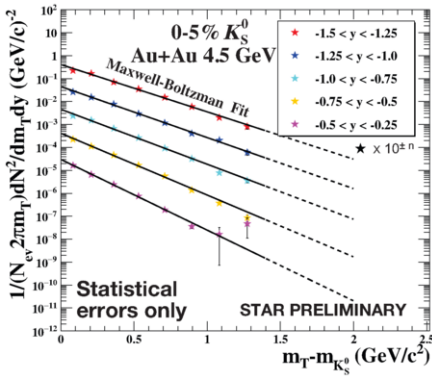


Fig. 3:  $K_S^0$  transverse mass spectra for central  $\sqrt{s_{NN}} = 4.5$  GeV Au+Au and 4.9 GeV Al+Au collisions.

In Fig. 3,  $m_T$  spectra of  $K_S^0$  from 4.5 GeV Au+Au and 4.9 GeV Al+Au collisions are shown for 0-5% centrality in several rapidity bins of width  $\Delta y = 0.25$ . These spectra are fitted using a Maxwell-Boltzmann functional form. The spectra and fits for  $\Lambda$  (not shown) are found to be similar to Fig. 3 in their range and fit quality. Kaon rapidity densities,  $dN/dy$ , are extracted from the fits and compared with AGS results in Fig. 4 (left). The STAR  $K_S^0$   $dN/dy$  falls between E917  $K^+$  and  $K^-$  results and are slightly lower than  $(K^+ + K^-)/2$  from E917 [6]. The amplitudes and widths of the current  $K_S^0$   $dN/dy$  indicate good agreement between STAR FXT and published AGS results.

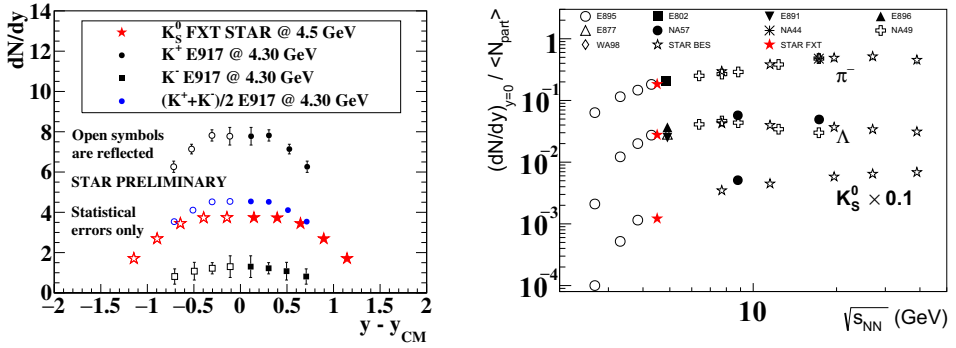


Fig. 4: (Left) Kaon  $dN/dy$  for Au+Au at or near 4.5 GeV. (Right) Peak  $dN/dy$  vs. beam energy for  $\pi^-$ ,  $\Lambda$  and  $K_S^0$ .

In Fig. 4 (right),  $dN/dy$  distributions of  $\pi^-$ ,  $K_S^0$  and  $\Lambda$  obtained in Au+Au collisions at  $\sqrt{s_{NN}} = 4.5$  GeV near mid-rapidity, normalized by mean participant number  $N_{part}$ , are compared across various experiments as a function of beam energy. The current STAR FXT rapidity densities are consistent with the trends defined by prior measurements at higher and lower energies [7–16].

## 2.2. STAR FXT Flow Results

The left panel of Fig. 5 presents proton  $v_1(y)$  from 4.5 GeV Au+Au (red markers). E895 data for 4.3 GeV Au+Au at similar  $p_T$  and centrality are also plotted [17]. STAR FXT results are consistent with E895 but have much smaller statistical errors. In the right panel of Fig. 5, proton  $v_1(y)$  is plotted in narrow intervals of  $p_T$ . The proton acceptance in FXT mode at this beam energy, with the standard selection of  $0.4 < p_T < 2.0$  GeV/c used in prior studies [3, 4, 17], was found to have negligible impact on proton  $v_1(y)$  down to  $y - y_{cm} = 0$ .

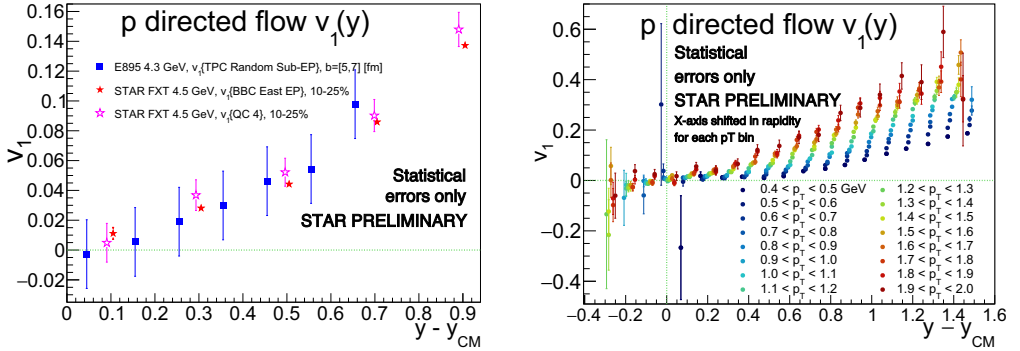


Fig. 5: (Left) proton  $v_1(y)$  for FXT 4.5 GeV Au+Au, based on two separate methods [18, 19], and compared with E895 [17]. (Right) The same FXT proton  $v_1(y)$  in narrow intervals of transverse momentum.

FXT directed flow slope  $dv_1/dy$  near midrapidity for protons,  $\pi^\pm$ ,  $K_S^0$  and  $\Lambda$  at 4.5 GeV are compared in the left panel of Fig. 6 with STAR collider-mode results [3, 4] and with E895 [17]. The STAR FXT  $\pi^\pm$  and  $K_S^0$  measurements continue the trend of negative  $dv_1/dy$  for mesons observed at 7.7 GeV and above. The STAR FXT  $dv_1/dy$  for protons at 4.5 GeV is in good agreement with E895 at 4.3 GeV, and both are consistent with a smooth interpolation between higher and lower energies.

Figure 7 presents FXT 4.5 GeV Au+Au proton and pion elliptic flow  $v_2(p_T)$ . The observed mass ordering of proton and pion  $v_2$  resembles measurements at higher energies [20–22].

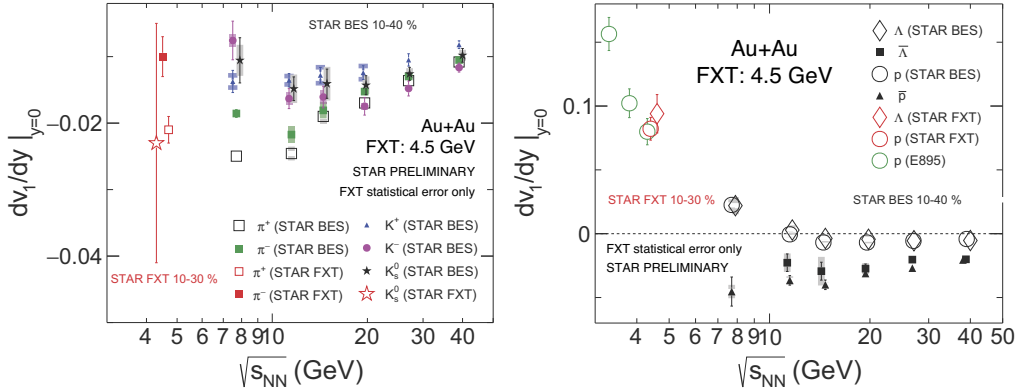


Fig. 6: Mid-rapidity directed flow slope versus beam energy for mesons (left panel) and for baryons (right panel).

### 3. Future FXT Plans

In future runs, STAR will have multiple detector upgrades [23], including the Inner Time Projection Chamber (iTPC), Endcap Time Of Flight (eTOF) and Event Plane Detector (EPD). The EPD is already in operation. These upgrades extend the detector acceptance, and improve the particle identification capability, event plane resolution and centrality determination. Running in FXT mode for two days at each of several energies between  $\sqrt{s_{NN}} = 3.0$  GeV and 7.7 GeV will permit acquisition of  $\sim 100$  million events per energy and will extend the reach of baryon chemical potential to about 720 MeV. The highest FXT energy will overlap with the lowest collider-mode energy, allowing detailed cross checks.

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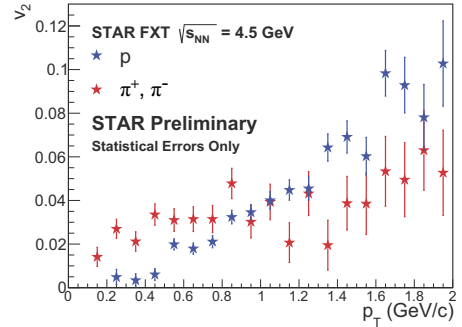


Fig. 7: 4.5 GeV Au+Au  $v_2(p_T)$  for  $\pi^\pm$  and protons.