

Measurement of D^0 Azimuthal Anisotropy at Midrapidity in Au + Au Collisions at $\sqrt{s_{NN}} = 200$ GeV

L. Adamczyk,¹ J. K. Adkins,¹⁹ G. Agakishiev,¹⁷ M. M. Aggarwal,³¹ Z. Ahammed,⁵⁰ N. N. Ajitanand,⁴⁰ I. Alekseev,^{15,26} D. M. Anderson,⁴² R. Aoyama,⁴⁶ A. Aparin,¹⁷ D. Arkhipkin,³ E. C. Aschenauer,³ M. U. Ashraf,⁴⁵ A. Attri,³¹ G. S. Averichev,¹⁷ X. Bai,⁷ V. Bairathi,²⁷ A. Behera,⁴⁰ R. Bellwied,⁴⁴ A. Bhasin,¹⁶ A. K. Bhati,³¹ P. Bhattarai,⁴³ J. Bielcik,¹⁰ J. Bielcikova,¹¹ L. C. Bland,³ I. G. Bordyuzhin,¹⁵ J. Bouchet,¹⁸ J. D. Brandenburg,³⁶ A. V. Brandin,²⁶ D. Brown,²³ I. Bunzarov,¹⁷ J. Butterworth,³⁶ H. Caines,⁵⁴ M. Calderón de la Barca Sánchez,⁵ J. M. Campbell,²⁹ D. Cebra,⁵ I. Chakaberia,³ P. Chaloupka,¹⁰ Z. Chang,⁴² N. Chankova-Bunzarova,¹⁷ A. Chatterjee,⁵⁰ S. Chattopadhyay,⁵⁰ X. Chen,³⁷ J. H. Chen,³⁹ X. Chen,²¹ J. Cheng,⁴⁵ M. Cherney,⁹ W. Christie,³ G. Contin,²² H. J. Crawford,⁴ S. Das,⁷ L. C. De Silva,⁹ R. R. Debbe,³ T. G. Dedovich,¹⁷ J. Deng,³⁸ A. A. Derevschikov,³³ L. Didenko,³ C. Dilks,³² X. Dong,²² J. L. Drachenberg,²⁰ J. E. Draper,⁵ L. E. Dunkelberger,⁶ J. C. Dunlop,³ L. G. Efimov,¹⁷ N. Elsey,⁵² J. Engelage,⁴ G. Eppley,³⁶ R. Esha,⁶ S. Esumi,⁴⁶ O. Evdokimov,⁸ J. Ewigleben,²³ O. Eyser,³ R. Fatemi,¹⁹ S. Fazio,³ P. Federic,¹¹ P. Federicova,¹⁰ J. Fedorisin,¹⁷ Z. Feng,⁷ P. Filip,¹⁷ E. Finch,⁴⁷ Y. Fisyak,³ C. E. Flores,⁵ L. Fulek,¹ C. A. Gagliardi,⁴² D. Garand,³⁴ F. Geurts,³⁶ A. Gibson,⁴⁹ M. Girard,⁵¹ L. Greiner,²² D. Grosnick,⁴⁹ D. S. Gunarathne,⁴¹ Y. Guo,¹⁸ A. Gupta,¹⁶ S. Gupta,¹⁶ W. Guryn,³ A. I. Hamad,¹⁸ A. Hamed,⁴² A. Harlanderova,¹⁰ J. W. Harris,⁵⁴ L. He,³⁴ S. Heppelmann,³² S. Heppelmann,⁵ A. Hirsch,³⁴ G. W. Hoffmann,⁴³ S. Horvat,⁵⁴ T. Huang,²⁸ B. Huang,⁸ X. Huang,⁴⁵ H. Z. Huang,⁶ T. J. Humanic,²⁹ P. Huo,⁴⁰ G. Igo,⁶ W. W. Jacobs,¹⁴ A. Jentsch,⁴³ J. Jia,^{3,40} K. Jiang,³⁷ S. Jowzaee,⁵² E. G. Judd,⁴ S. Kabana,¹⁸ D. Kalinkin,¹⁴ K. Kang,⁴⁵ K. Kauder,⁵² H. W. Ke,³ D. Keane,¹⁸ A. Kechechyan,¹⁷ Z. Khan,⁸ D. P. Kikoła,⁵¹ I. Kisel,¹² A. Kisel,⁵¹ L. Kochenda,²⁶ M. Kocmanek,¹¹ T. Kollegger,¹² L. K. Kosarzewski,⁵¹ A. F. Kraishan,⁴¹ P. Kravtsov,²⁶ K. Krueger,² N. Kulathunga,⁴⁴ L. Kumar,³¹ J. Kvapil,¹⁰ J. H. Kwasizur,¹⁴ R. Lacey,⁴⁰ J. M. Landgraf,³ K. D. Landry,⁶ J. Lauret,³ A. Lebedev,³ R. Lednicky,¹⁷ J. H. Lee,³ X. Li,³⁷ C. Li,³⁷ W. Li,³⁹ Y. Li,⁴⁵ J. Lidrych,¹⁰ T. Lin,¹⁴ M. A. Lisa,²⁹ H. Liu,¹⁴ P. Liu,⁴⁰ Y. Liu,⁴² F. Liu,⁷ T. Ljubicic,³ W. J. Llope,⁵² M. Lomnitz,²² R. S. Longacre,³ S. Luo,⁸ X. Luo,⁷ G. L. Ma,³⁹ L. Ma,³⁹ Y. G. Ma,³⁹ R. Ma,³ N. Magdy,⁴⁰ R. Majka,⁵⁴ D. Mallick,²⁷ S. Margetis,¹⁸ C. Markert,⁴³ H. S. Matis,²² K. Meehan,⁵ J. C. Mei,³⁸ Z. W. Miller,⁸ N. G. Minaev,³³ S. Mioduszewski,⁴² D. Mishra,²⁷ S. Mizuno,²² B. Mohanty,²⁷ M. M. Mondal,¹³ D. A. Morozov,³³ M. K. Mustafa,²² Md. Nasim,⁶ T. K. Nayak,⁵⁰ J. M. Nelson,⁴ M. Nie,³⁹ G. Nigmatkulov,²⁶ T. Niida,⁵² L. V. Nogach,³³ T. Nonaka,⁴⁶ S. B. Nurushev,³³ G. Odyniec,²² A. Ogawa,³ K. Oh,³⁵ V. A. Okorokov,²⁶ D. Olivitt, Jr.,⁴¹ B. S. Page,³ R. Pak,³ Y. Pandit,⁸ Y. Panebratsev,¹⁷ B. Pawlik,³⁰ H. Pei,⁷ C. Perkins,⁴ P. Pile,³ J. Pluta,⁵¹ K. Poniatowska,⁵¹ J. Porter,²² M. Posik,⁴¹ A. M. Poskanzer,²² N. K. Pruthi,³¹ M. Przybycien,¹ J. Putschke,⁵² H. Qiu,³⁴ A. Quintero,⁴¹ S. Ramachandran,¹⁹ R. L. Ray,⁴³ R. Reed,²³ M. J. Rehbein,⁹ H. G. Ritter,²² J. B. Roberts,³⁶ O. V. Rogachevskiy,¹⁷ J. L. Romero,⁵ J. D. Roth,⁹ L. Ruan,³ J. Rusnak,¹¹ O. Rusnakova,¹⁰ N. R. Sahoo,⁴² P. K. Sahu,¹³ S. Salur,²² J. Sandweiss,⁵⁴ M. Saur,¹¹ J. Schambach,⁴³ A. M. Schmah,²² W. B. Schmidke,³ N. Schmitz,²⁴ B. R. Schweid,⁴⁰ J. Seger,⁹ M. Sergeeva,⁶ P. Seyboth,²⁴ N. Shah,³⁹ E. Shahaliev,¹⁷ P. V. Shanmuganathan,²³ M. Shao,³⁷ A. Sharma,¹⁶ M. K. Sharma,¹⁶ W. Q. Shen,³⁹ Z. Shi,²² S. S. Shi,⁷ Q. Y. Shou,³⁹ E. P. Sichtermann,²² R. Sikora,¹ M. Simko,¹¹ S. Singha,¹⁸ M. J. Skoby,¹⁴ N. Smirnov,⁵⁴ D. Smirnov,³ W. Solyst,¹⁴ L. Song,⁴⁴ P. Sorensen,³ H. M. Spinka,² B. Srivastava,³⁴ T. D. S. Stanislaus,⁴⁹ M. Strikhanov,²⁶ B. Stringfellow,³⁴ T. Sugiura,⁴⁶ M. Sumera,¹¹ B. Summa,³² Y. Sun,³⁷ X. M. Sun,⁷ X. Sun,⁷ B. Surrow,⁴¹ D. N. Svirida,¹⁵ M. A. Szelezniak,²² A. H. Tang,³ Z. Tang,³⁷ A. Taranenko,²⁶ T. Tarnowsky,²⁵ A. Tawfik,⁵³ J. Thäder,²² J. H. Thomas,²² A. R. Timmins,⁴⁴ D. Tlusty,³⁶ T. Todoroki,³ M. Tokarev,¹⁷ S. Trentalange,⁶ R. E. Tribble,⁴² P. Tribedy,³ S. K. Tripathy,¹³ B. A. Trzeciak,¹⁰ O. D. Tsai,⁶ T. Ullrich,³ D. G. Underwood,² I. Upsal,²⁹ G. Van Buren,³ G. van Nieuwenhuizen,³ A. N. Vasiliev,³³ F. Videbæk,³ S. Vokal,¹⁷ S. A. Voloshin,⁵² A. Vossen,¹⁴ G. Wang,⁶ Y. Wang,⁷ F. Wang,³⁴ Y. Wang,⁴⁵ J. C. Webb,³ G. Webb,³ L. Wen,⁶ G. D. Westfall,²⁵ H. Wieman,²² S. W. Wissink,¹⁴ R. Witt,⁴⁸ Y. Wu,¹⁸ Z. G. Xiao,⁴⁵ W. Xie,³⁴ G. Xie,³⁷ J. Xu,⁷ N. Xu,²² Q. H. Xu,³⁸ Y. F. Xu,³⁹ Z. Xu,³ Y. Yang,²⁸ Q. Yang,³⁷ C. Yang,³⁸ S. Yang,³ Z. Ye,⁸ Z. Ye,⁸ L. Yi,⁵⁴ K. Yip,³ I.-K. Yoo,³⁵ N. Yu,⁷ H. Zbroszczyk,⁵¹ W. Zha,³⁷ Z. Zhang,³⁹ X. P. Zhang,⁴⁵ J. B. Zhang,⁷ S. Zhang,³⁷ J. Zhang,²¹ Y. Zhang,³⁷ J. Zhang,²² S. Zhang,³⁹ J. Zhao,³⁴ C. Zhong,³⁹ L. Zhou,³⁷ C. Zhou,³⁹ X. Zhu,⁴⁵ Z. Zhu,³⁸ and M. Zyzak¹²

(STAR Collaboration)

¹AGH University of Science and Technology, FPACS, Cracow 30-059, Poland²Argonne National Laboratory, Argonne, Illinois 60439

- ³Brookhaven National Laboratory, Upton, New York 11973
⁴University of California, Berkeley, California 94720
⁵University of California, Davis, California 95616
⁶University of California, Los Angeles, California 90095
⁷Central China Normal University, Wuhan, Hubei 430079
⁸University of Illinois at Chicago, Chicago, Illinois 60607
⁹Creighton University, Omaha, Nebraska 68178
¹⁰Czech Technical University in Prague, FNSPE, Prague 115 19, Czech Republic
¹¹Nuclear Physics Institute AS CR, 250 68 Prague, Czech Republic
¹²Frankfurt Institute for Advanced Studies FIAS, Frankfurt 60438, Germany
¹³Institute of Physics, Bhubaneswar 751005, India
¹⁴Indiana University, Bloomington, Indiana 47408
¹⁵Alikhanov Institute for Theoretical and Experimental Physics, Moscow 117218, Russia
¹⁶University of Jammu, Jammu 180001, India
¹⁷Joint Institute for Nuclear Research, Dubna 141 980, Russia
¹⁸Kent State University, Kent, Ohio 44242
¹⁹University of Kentucky, Lexington, Kentucky 40506-0055
²⁰Lamar University, Physics Department, Beaumont, Texas 77710
²¹Institute of Modern Physics, Chinese Academy of Sciences, Lanzhou, Gansu 730000
²²Lawrence Berkeley National Laboratory, Berkeley, California 94720
²³Lehigh University, Bethlehem, Pennsylvania 18015
²⁴Max-Planck-Institut für Physik, Munich 80805, Germany
²⁵Michigan State University, East Lansing, Michigan 48824
²⁶National Research Nuclear University MEPhI, Moscow 115409, Russia
²⁷National Institute of Science Education and Research, HBNI, Jatni 752050, India
²⁸National Cheng Kung University, Tainan 70101
²⁹Ohio State University, Columbus, Ohio 43210
³⁰Institute of Nuclear Physics PAN, Cracow 31-342, Poland
³¹Panjab University, Chandigarh 160014, India
³²Pennsylvania State University, University Park, Pennsylvania 16802
³³Institute of High Energy Physics, Protvino 142281, Russia
³⁴Purdue University, West Lafayette, Indiana 47907
³⁵Pusan National University, Pusan 46241, Korea
³⁶Rice University, Houston, Texas 77251
³⁷University of Science and Technology of China, Hefei, Anhui 230026
³⁸Shandong University, Jinan, Shandong 250100
³⁹Shanghai Institute of Applied Physics, Chinese Academy of Sciences, Shanghai 201800
⁴⁰State University of New York, Stony Brook, New York 11794
⁴¹Temple University, Philadelphia, Pennsylvania 19122
⁴²Texas A&M University, College Station, Texas 77843
⁴³University of Texas, Austin, Texas 78712
⁴⁴University of Houston, Houston, Texas 77204
⁴⁵Tsinghua University, Beijing 100084
⁴⁶University of Tsukuba, Tsukuba, Ibaraki, Japan
⁴⁷Southern Connecticut State University, New Haven, Connecticut 06515
⁴⁸United States Naval Academy, Annapolis, Maryland 21402
⁴⁹Valparaiso University, Valparaiso, Indiana 46383
⁵⁰Variable Energy Cyclotron Centre, Kolkata 700064, India
⁵¹Warsaw University of Technology, Warsaw 00-661, Poland
⁵²Wayne State University, Detroit, Michigan 48201
⁵³World Laboratory for Cosmology and Particle Physics (WLCAPP), Cairo 11571, Egypt
⁵⁴Yale University, New Haven, Connecticut 06520

(Received 23 January 2017; revised manuscript received 31 March 2017; published 26 May 2017)

We report the first measurement of the elliptic anisotropy (v_2) of the charm meson D^0 at midrapidity ($|y| < 1$) in Au + Au collisions at $\sqrt{s_{NN}} = 200$ GeV. The measurement was conducted by the STAR experiment at RHIC utilizing a new high-resolution silicon tracker. The measured D^0 v_2 in 0%–80% centrality Au + Au collisions can be described by a viscous hydrodynamic calculation for a transverse momentum (p_T) of less than 4 GeV/c. The D^0 v_2 as a function of transverse kinetic energy ($m_T - m_0$, where $m_T = \sqrt{p_T^2 + m_0^2}$) is consistent with that of light mesons in 10%–40% centrality Au + Au

collisions. These results suggest that charm quarks have achieved local thermal equilibrium with the medium created in such collisions. Several theoretical models, with the temperature-dependent, dimensionless charm spatial diffusion coefficient ($2\pi TD_s$) in the range of ~ 2 – 12 , are able to simultaneously reproduce our $D^0 v_2$ result and our previously published results for the D^0 nuclear modification factor.

DOI: [10.1103/PhysRevLett.118.212301](https://doi.org/10.1103/PhysRevLett.118.212301)

Quantum chromodynamics (QCD) is a non-Abelian gauge theory which describes the strong interactions between quarks and gluons. Experiments at the Relativistic Heavy Ion Collider (RHIC) and the Large Hadron Collider (LHC) indicate that a novel form of QCD matter, consistent with a strongly coupled quark-gluon plasma (sQGP), is created in heavy-ion collisions at these energies [1–3]. A key piece of evidence for this new state of matter is the strong collective, anisotropic flow of produced light flavor particles, suggesting possibly hydrodynamic behavior of the strongly interacting matter during the collision [4].

Heavy quarks (charm and bottom) are predominantly created in the initial hard scatterings in a heavy-ion collision, and their propagation in the sQGP can be described as Brownian-like motion [5,6]. The sQGP properties can be accessed through experimental observables such as the nuclear modification factor (R_{AA}) [7], the ratio of the yield in heavy-ion collisions to the scaled yield in proton + proton ($p + p$) collisions, and the elliptic anisotropy (v_2) [8], the second Fourier coefficient of the particle yield with respect to the reaction plane (defined by the beam axis and the direction of the impact parameter between two colliding nuclei). Of these observables, the v_2 at a low transverse momentum (p_T), where light and strange flavor hadrons appear to behave hydrodynamically, is of particular interest, because it probes the properties of the bulk medium in the strongly coupled region and is less affected by the shadowing and Cronin effects [9].

Recent measurements at RHIC and the LHC show that high- p_T charm hadron yields are significantly suppressed in central heavy-ion collisions indicating strong charm-medium interactions [10–12]. The D -meson v_2 measured by ALICE [13] is comparable to that of light hadrons at the LHC. So far, charm quark flow at RHIC has only been inferred from measurements of semileptonic decays of charm and bottom hadrons [14,15]. However, a clear interpretation of lepton v_2 measurements suffers from an ambiguity in the lepton sources between charm and bottom decays and the decay kinematics. On the other hand, there has been significant progress in theoretical calculations for charm hadron v_2 in heavy-ion collisions [16–23]. A precise measurement of charm hadron v_2 over a wide momentum range is expected to provide valuable insights into the sQGP properties [9].

In this Letter, we report the first measurement of the D^0 anisotropy parameter v_2 at midrapidity ($|y| < 1$) at RHIC

by the STAR Collaboration using the newly completed heavy flavor tracker (HFT) [24,25]. The HFT is a high-resolution silicon detector system, which aims for the topological reconstruction of secondary decay vertices of open heavy flavor hadrons. It has three subdetectors: the silicon strip detector, the intermediate silicon tracker (IST), and the pixel (PXL) detector. In the 2014 Au + Au run at $\sqrt{s_{NN}} = 200$ GeV, ~ 1.1 billion minimum bias triggered events, selected by a coincidence signal between the east and west vertex position detectors (VPDs) [26] located at $4.4 < |\eta| < 4.9$ (η is the pseudorapidity), were recorded with the IST and the PXL. In this analysis, the reconstructed collision primary vertex (PV) is required to be less than 6 cm from the detector center along the beam axis to ensure good HFT acceptance. The collision centrality, the fraction of the total hadronic cross section, is defined using the measured charged track multiplicity at midrapidity and corrected for the online VPD triggering inefficiency using a Monte Carlo Glauber simulation [27].

D^0 and \bar{D}^0 mesons are reconstructed in the $K^\mp \pi^\pm$ decay channel, which has a short proper decay length ($c\tau \sim 123 \mu\text{m}$) [28]. Charged tracks are reconstructed by the time projection chamber (TPC) [29] together with the HFT in a 0.5 T uniform magnetic field. Tracks are required to have a minimum of 20 TPC hits (out of a maximum of 45), hits in all layers of PXL and IST subdetectors, $p_T > 0.6$ GeV/ c , and $|\eta| < 1$. To identify particle species, the ionization energy loss dE/dx measured by the TPC is required to be within 3 and 2 standard deviations from the expected values for π and K , respectively. The particle identification is extended by the time of flight (TOF) [30] detector up to $p_T \sim 1.6$ GeV/ c by requiring the $1/\beta$ (β is the particle velocity in units of the speed of light), calculated from the path length and the TOF, to be less than 3 standard deviations different from the expected value calculated using the π or K mass and the measured momentum.

Figure 1(a) shows the track pointing resolution to the collision vertex in the transverse plane (σ_{XY}) as a function of momentum (p) for identified particles in 0%–80% centrality Au + Au collisions at $\sqrt{s_{NN}} = 200$ GeV. The resolution is better than $55 \mu\text{m}$ for kaons with $p \geq 0.75$ GeV/ c . With two daughter tracks, a secondary decay vertex can be reconstructed as the middle point on the distance of the closest approach (DCA) between them. The primary background is due to fake pairs coming from random combinations of tracks which propagate directly from the collision point. The background can be

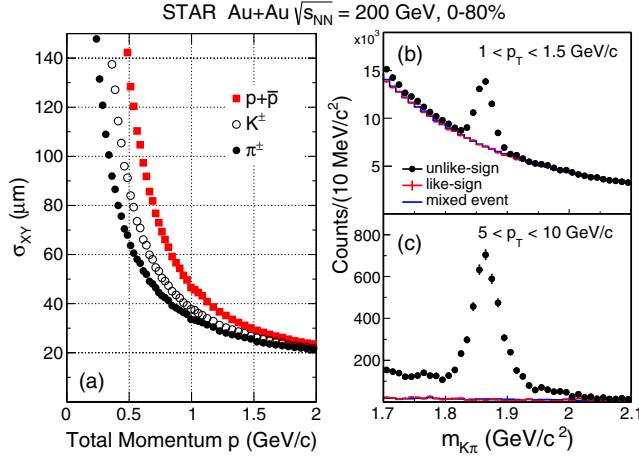


FIG. 1. Identified particle pointing resolution in the transverse plane as a function of particle momentum (a). Invariant mass spectra of $K\pi$ pairs for $1 < p_T < 1.5$ GeV/c (b) and $5 < p_T < 10$ GeV/c (c), respectively. The solid data points are the D^0 signal reconstructed with unlike-sign pairs. The red crosses and the blue lines show the like-sign and mixed event background distributions, respectively.

significantly reduced by applying cuts on five variables: the decay length (the distance between the decay vertex and the PV), the DCA between the two daughters, the DCA between the reconstructed D^0 track and the PV, the DCA between the π track and the PV, and the DCA between the K track and the PV. The cuts on these variables are optimized using the Ttoolkit for Multivariate Data Analysis package [31]. Their optimization was pursued separately in each D^0 candidate p_T bin in order to have the greatest signal significance.

Figures 1(b) and 1(c) show the invariant mass spectra of $K\pi$ pairs after applying these cuts for two p_T bins. Comparing these mass spectra with the previous D^0 study [10], the signal significance is markedly improved due to the background rejection using the geometric cuts enabled by the HFT ($\sim 220\sigma$ vs $\sim 13\sigma$ per billion events). The combinatorial background is estimated with like-sign $K\pi$ pairs and the mixed event unlike-sign technique in which K and π with opposite charge signs from different events are paired. The mixed event distributions are normalized to the like-sign distributions in the mass range of 1.7–2.1 GeV/ c^2 . The remaining contributions to the background are expected to come from the correlated sources, e.g., $K\pi$ pairs from jet fragments or multiprong decays of heavy flavor mesons.

Two different methods are employed to calculate v_2 : the event plane method [8] and the correlation method [32,33]. In the event plane method, a second-order event plane angle Ψ_2 is reconstructed from TPC tracks excluding decay products of D^0 mesons and after correcting for the azimuthal nonuniformity in the detector efficiency [8]. To suppress nonflow effects (correlations not connected to the event plane, such as resonance decays and jet

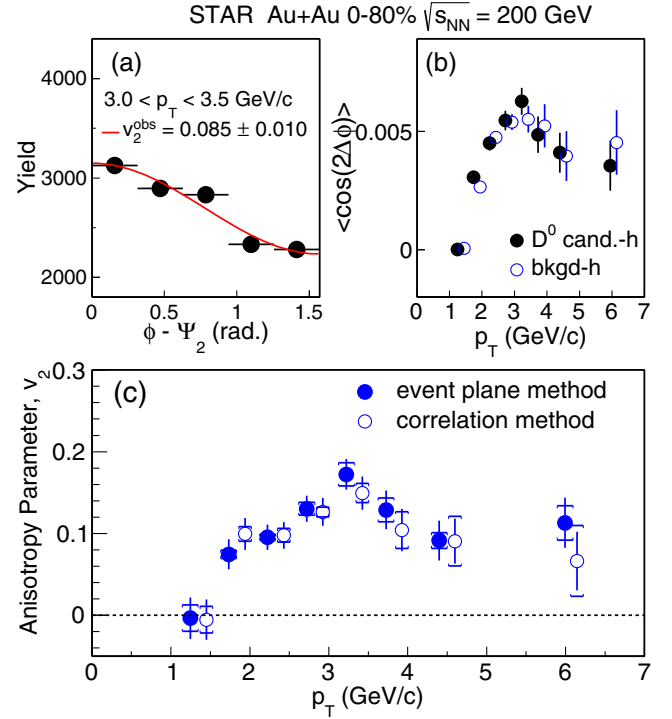


FIG. 2. (a) D^0 yield as a function of $\phi - \Psi_2$ fit to $A\{1 + 2v_2 \cos[2(\phi - \Psi_2)]\}$, for $3 < p_T < 3.5$ GeV/c. (b) Correlations $\langle \cos(2\Delta\phi) \rangle$ between the D^0 candidate or background and charged particles, as a function of p_T . (c) v_2 as a function of p_T for D^0 calculated with the event plane and correlation methods. The data shown in all three panels are for 0%–80% centrality Au + Au collisions at $\sqrt{s_{NN}} = 200$ GeV. The vertical bars and the brackets represent the statistical and systematic uncertainties, respectively. The estimated nonflow contribution is not shown in this plot but is common to both methods. In (a) and (b), only statistical uncertainties are shown as vertical bars (not visible if they are smaller than marker sizes). In (b) and (c), the open points are shifted along the x axis for clarity.

correlations), only particles from the opposite η hemisphere of the reconstructed D^0 and outside of an additional η gap of $|\Delta\eta| > 0.05$ are used in the event plane reconstruction. The D^0 yields are measured in azimuthal bins relative to the event plane azimuth ($\phi - \Psi_2$). The yields are weighted by $1/(\epsilon \times R)$, where ϵ is the D^0 reconstruction efficiency \times acceptance and R the event plane angle resolution [8] for each centrality interval [34]. In each $\phi - \Psi_2$ bin, the mixed event background, scaled to the like-sign background, is subtracted from the unlike-sign distribution. The D^0 yield is obtained via the sideband method by subtracting the scaled counts in two invariant mass ranges around the signal (1.71–1.80 and 1.93–2.02 GeV/ c^2) from the counts in the signal region (1.82–1.91 GeV/ c^2) [35]. A fit method using a Gaussian function for the D^0 signal plus a first-order polynomial function for the background is also used to estimate the systematic uncertainty on the raw yield extraction. Figure 2(a) shows an example of the weighted

D^0 yield as a function of $\phi - \Psi_2$. The observed v_2 is then obtained by fitting with a functional form $A\{1 + 2v_2 \cos[2(\phi - \Psi_2)]\}$, where A is a normalization parameter. Finally, the true v_2 is obtained by scaling the observed v_2 with $\langle 1/R \rangle$ to correct for the event plane angle resolution [34].

In the correlation method [32,33], v_2 is calculated for D^0 candidates and the background, separately. For example, the D^0 candidate-hadron azimuthal cumulant $V_2^{\text{cand-h}} \equiv \langle \cos(2\phi_{\text{cand}} - 2\phi_h) \rangle$, shown as a function of p_T as solid markers in Fig. 2(b), is calculated by the Q -cumulant method where ϕ_{cand} and ϕ_h are azimuthal angles for D^0 candidates and charged hadrons, respectively [33]. The average is taken over all events and all particles. Neglecting nonflow contributions, the following factorization can be assumed to obtain the D^0 v_2 : $V_2^{\text{cand-h}} = v_2^{\text{cand}} v_2^h$. Here, v_2^h can be obtained from hadron-hadron correlations via $V_2^{h-h} = v_2^h v_2^h$. The same η gap as in the event plane method was chosen for the correlation analysis. The D^0 background v_2 is calculated similarly, with the background represented by the average of the like-sign $K\pi$ pairs in the D^0 mass window ($\pm 3\sigma$, where σ is the signal width) and sidebands ($4-9\sigma$ away from the D^0 peak, both like-sign and unlike-sign $K\pi$ pairs). The background-hadron cumulant is also shown in Fig. 2(b) as open circles. The D^0 v_2 is obtained from the candidate and background v_2 and their respective yields (N_{cand} , N_{bg}) by $v_2 = (N_{\text{cand}} v_2^{\text{cand}} - N_{\text{bg}} v_2^{\text{bg}}) / (N_{\text{cand}} - N_{\text{bg}})$.

The systematic uncertainty is estimated by comparing v_2 obtained from the following different methods: (a) the fit vs sideband methods, (b) varying invariant mass ranges for the fit and for the sidebands, and (c) varying geometric cuts so that the efficiency changes by $\pm 50\%$ with respect to the nominal value. These three different sources are varied independently to form multiple combinations. We then take the maximum difference from these combinations and divide by $\sqrt{12}$ as one standard deviation of the systematic uncertainty. The feed-down contribution from B -meson decays to our measured D^0 yield is estimated to be less than 4%. Compared to other systematic uncertainties, this contribution is negligible even in the extreme case that B -meson v_2 is 0.

Figure 2(c) shows the result of the D^0 v_2 in 0%–80% centrality Au + Au events as a function of p_T . The results from the event plane and correlation methods are consistent with each other within uncertainties. For further discussion in this Letter, we use v_2 from the event plane method only, which has been widely used in previous STAR-identified particle v_2 measurements [36,37].

The residual nonflow contribution is estimated by scaling the D^0 -hadron correlation (with the same η gap used in the analysis) in $p + p$ collisions, where only the nonflow effects are present, by the average v_2 (\bar{v}_2) and multiplicity (M) of charged hadrons used for event plane reconstruction or D^0 -hadron correlations in Au + Au collisions. Thus, the

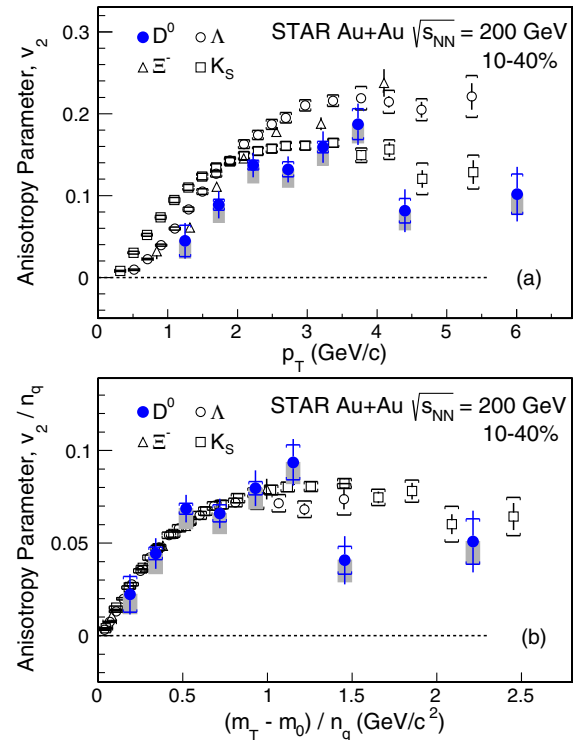


FIG. 3. (a) v_2 as a function of p_T and (b) v_2/n_q as a function of $(m_T - m_0)/n_q$ for D^0 in 10%–40% centrality Au + Au collisions compared with K_S^0 , Λ , and Ξ^- [36]. The vertical bars and brackets represent statistical and systematic uncertainties, respectively, and the gray bands represent the estimated nonflow contribution.

nonflow contribution is estimated to be $\langle \sum_i \cos 2(\phi_{D^0} - \phi_i) \rangle / M \bar{v}_2$ [38], where ϕ_{D^0} and ϕ_i are the azimuthal angles for the D^0 and hadron, respectively. The \sum_i is done for charged tracks in the same event, and $\langle \rangle$ is an average over all events. The D^0 -hadron correlation in $p + p$ collisions is deduced from D^* -hadron correlations measured with data taken by STAR in the year 2012 for $p_T > 3$ GeV/c and from a PYTHIA simulation for $p_T < 3$ GeV/c. The correlations in $p + p$ collisions were used as a conservative estimate, since the correlation may be suppressed in Au + Au collisions due to the hot medium effect. The estimated nonflow contribution is shown separately (gray bands) along with the systematic and statistical uncertainties in Figs. 3 and 4.

For a cross-check, we performed a MC simulation using the measured D^0 v_2 to calculate the single electron v_2 and compare to previous RHIC measurements [14,15]. Both the PHENIX and STAR measurements are compatible with the calculated electron v_2 at $p_T < 3$ GeV/c, where the charm hadron contribution dominates [40–42]. At a higher p_T region, where the bottom contribution is sizable, the large uncertainty in the measurement of v_2 of single electrons does not allow for a reasonable extraction of v_2 for B mesons.

Figure 3 compares the measured D^0 v_2 from the event plane method in 10%–40% centrality bin with v_2 of K_S^0 , Λ , and Ξ^- [36]. The comparison between D^0 and light hadrons

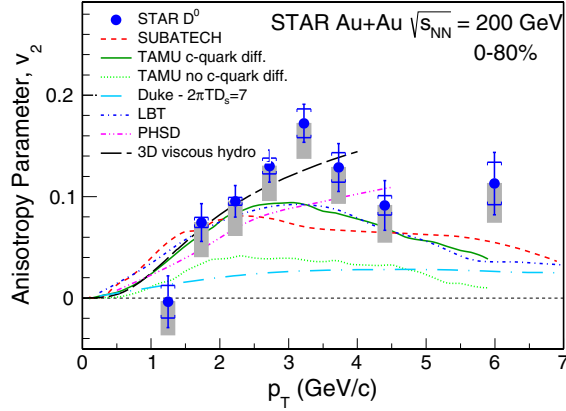


FIG. 4. v_2 as a function of p_T for D^0 in 0%–80% centrality Au + Au collisions compared with model calculations [16–21,39].

needs to be done in a narrow centrality bin to avoid the bias caused by the fact that the D^0 yield scales with a number of binary collisions while the yield of light hadrons scales approximately with a number of the participants [43]. Figure 3(a) shows v_2 as a function of p_T where a clear mass ordering for $p_T < 2$ GeV/c including D^0 mesons is observed. For $p_T > 2$ GeV/c, the D^0 meson v_2 follows that of other light mesons indicating a significant charm quark flow at RHIC [36,37,44]. Recent ALICE measurements show that the D^0 v_2 is comparable to that of charged hadrons in 0%–50% Pb + Pb collisions at $\sqrt{s_{NN}} = 2.76$ TeV [13] suggesting a sizable charm flow at the LHC. Figure 3(b) shows v_2/n_q as a function of scaled transverse kinetic energy $(m_T - m_0)/n_q$, where n_q is the number of constituent quarks in the hadron, m_0 its mass, and $m_T = \sqrt{p_T^2 + m_0^2}$. We find that the D^0 v_2 falls into the same universal trend as all other light hadrons [45], in particular, for $(m_T - m_0)/n_q < 1$ GeV/c². This suggests that charm quarks have gained significant flow through interactions with the sQGP medium in 10%–40% Au + Au collisions at $\sqrt{s_{NN}} = 200$ GeV.

The heavy quark-medium interaction is often characterized by a spatial diffusion coefficient D_s , or a dimensionless coefficient $2\pi TD_s$, where T is the medium temperature [5]. In Fig. 4, the measured D^0 v_2 in 0%–80% centrality collisions is compared with several model calculations [16–21,39]. The Duke, LBT, PHSD, and SUBATECH models and the TAMU model with charm quark diffusion are able to describe our previously published D^0 R_{AA} result [10,16,21]. Compared to the v_2 measurement, the TAMU model with no charm quark diffusion does not reproduce the data, while the same model with charm quark diffusion turned on describes the data better [20]. A 3D viscous event-by-event hydrodynamic simulation with $\eta/s = 0.12$ using the AMPT initial condition and tuned to describe v_2 for light hadrons predicts D^0 v_2 that is consistent with our data for $p_T < 4$ GeV/c [39]. This suggests that charm quarks have achieved thermal equilibrium in these collisions.

TABLE I. D^0 v_2 in 0%–80% centrality Au + Au collisions compared with model calculations, quantified by χ^2/NDF and the p value. $2\pi TD_s$ values quoted are in the range of T_c to $2T_c$. χ^2/NDF is calculated in the p_T range wherever the model calculation is available.

Compare with	$2\pi TD_s$	χ^2/NDF	p value
SUBATECH [17]	2–4	15.2/8	0.06
TAMU c quark diffusion [20]	5–12	10.0/8	0.26
TAMU no c quark diffusion [20]	...	29.5/8	2×10^{-4}
Duke [19]	7	35.7/8	2×10^{-5}
LBT [21]	3–6	11.1/8	0.19
PHSD [16]	5–12	8.7/7	0.28
3D viscous hydro [39]	...	3.6/6	0.73

We performed a statistical significance test for the consistency between our data and each model quantified by χ^2/NDF and the p value listed in Table I. One can observe that the Duke model and TAMU model with no charm quark diffusion are inconsistent with our v_2 data, while other models describe the v_2 data in the measured p_T region. These models that can describe both the R_{AA} and v_2 data include the temperature-dependent charm diffusion coefficient $2\pi TD_s$ in the range of ~ 2 –12. $2\pi TD_s$ predicted by lattice QCD calculations fall in the same range [46,47]. In addition to the different treatments of the charm-medium interactions, there are also various differences among these models, e.g., the initial state, the space-time description of the QGP evolution, the hadronization, and the interactions in the hadronic matter. More coherent model treatments of these aspects are needed in order to better interpret the information about charm-medium interaction and provide a better constraint on $2\pi TD_s$ using our D^0 v_2 measurement.

In summary, the D^0 v_2 in Au + Au collisions at $\sqrt{s_{NN}} = 200$ GeV has been measured with the STAR detector using the heavy flavor tracker, a newly installed high-resolution silicon detector. The measured D^0 v_2 follows the mass ordering at low p_T observed earlier. The v_2/n_q of D^0 is consistent with that of other hadrons at $(m_T - m_0)/n_q < 1$ GeV/c² in 10%–40% centrality collisions. A 3D viscous hydrodynamic model describes the D^0 v_2 for $p_T < 4$ GeV/c. Our results suggest that charm quarks exhibit the same strong collective behavior as the light hadrons and may be close to thermal equilibrium in Au + Au collisions at $\sqrt{s_{NN}} = 200$ GeV. Several theoretical calculations with temperature-dependent, dimensionless charm quark spatial diffusion coefficients ($2\pi TD_s$) in the range of ~ 2 –12 can simultaneously reproduce our D^0 v_2 result as well as the previously published STAR measurement of the D^0 nuclear modification factor. The charm quark diffusion coefficients from lattice QCD calculations are consistent with the same range [46,47].

We thank the RHIC Operations Group and RCF at BNL, the NERSC Center at LBNL, and the Open Science Grid

consortium for providing resources and support. This work was supported in part by the Office of Nuclear Physics within the U.S. DOE Office of Science, the U.S. National Science Foundation, the Ministry of Education and Science of the Russian Federation, National Natural Science Foundation of China, Chinese Academy of Science, the Ministry of Science and Technology of China (973 Programme No. 2014CB845400) and the Chinese Ministry of Education, the National Research Foundation of Korea, GA and MSMT of the Czech Republic, Department of Atomic Energy and Department of Science and Technology of the Government of India; the National Science Centre of Poland, National Research Foundation, the Ministry of Science, Education and Sports of the Republic of Croatia, RosAtom of Russia and German Bundesministerium für Bildung, Wissenschaft, Forschung and Technologie (BMBF), and the Helmholtz Association.

-
- [1] J. Adams *et al.* (STAR Collaboration), *Nucl. Phys.* **A757**, 102 (2005).
- [2] K. Adcox *et al.* (PHENIX Collaboration), *Nucl. Phys.* **A757**, 184 (2005).
- [3] B. Müller, J. Schukraft, and B. Wyslouch, *Annu. Rev. Nucl. Part. Sci.* **62**, 361 (2012).
- [4] C. Gale, S. Jeon, B. Schenke, P. Tribedy, and R. Venugopalan, *Phys. Rev. Lett.* **110**, 012302 (2013).
- [5] G. D. Moore and D. Teaney, *Phys. Rev. C* **71**, 064904 (2005).
- [6] R. Rapp and H. van Hees, [arXiv:0803.0901](https://arxiv.org/abs/0803.0901).
- [7] X.-N. Wang and M. Gyulassy, *Phys. Rev. Lett.* **68**, 1480 (1992).
- [8] A. M. Poskanzer and S. A. Voloshin, *Phys. Rev. C* **58**, 1671 (1998).
- [9] A. Andronic *et al.*, *Eur. Phys. J. C* **76**, 107 (2016).
- [10] L. Adamczyk *et al.* (STAR Collaboration), *Phys. Rev. Lett.* **113**, 142301 (2014).
- [11] B. Abelev *et al.* (ALICE Collaboration), *J. High Energy Phys.* **09** (2012) 112.
- [12] J. Adam *et al.* (ALICE Collaboration), *J. High Energy Phys.* **03** (2016) 081.
- [13] B. Abelev *et al.* (ALICE Collaboration), *Phys. Rev. Lett.* **111**, 102301 (2013); *Phys. Rev. C* **90**, 034904 (2014).
- [14] A. Adare *et al.* (PHENIX Collaboration), *Phys. Rev. Lett.* **98**, 172301 (2007); *Phys. Rev. C* **84**, 044905 (2011).
- [15] L. Adamczyk *et al.* (STAR Collaboration), *Phys. Rev. C* **95**, 034907 (2017).
- [16] H. Berrehrh, P. B. Gossiaux, J. Aichelin, W. Cassing, J. M. Torres-Rincon, and E. Bratkovskaya, *Phys. Rev. C* **90**, 051901 (2014); T. Song, H. Berrehrh, D. Cebreira, J. M. Torres-Rincon, L. Tolos, W. Cassing, and E. Bratkovskaya, *Phys. Rev. C* **92**, 014910 (2015).
- [17] V. Ozvenchuk, J. M. Torres-Rincon, P. B. Gossiaux, J. Aichelin, and L. Tolos, *Phys. Rev. C* **90**, 054909 (2014); M. Nahrgang, J. Aichelin, S. Bass, P. B. Gossiaux, and K. Werner, *Phys. Rev. C* **91**, 014904 (2015); (private communication).
- [18] W. M. Alberico, A. Beraudo, A. De Pace, A. Molinari, M. Monteno, M. Nardi, and F. Prino, *Eur. Phys. J. C* **71**, 1666 (2011); A. Beraudo, A. De Pace, M. Monteno, M. Nardi, and F. Prino, *Eur. Phys. J. C* **75**, 121 (2015).
- [19] S. Cao, G. Y. Qin, and S. A. Bass, *Phys. Rev. C* **92**, 024907 (2015); (private communication).
- [20] M. He, R. J. Fries, and R. Rapp, *Phys. Rev. C* **86**, 014903 (2012); *Phys. Rev. Lett.* **110**, 112301 (2013); (private communication).
- [21] S. Cao, T. Luo, G.-Y. Qin, and X.-N. Wang, *Phys. Rev. C* **94**, 014909 (2016).
- [22] J. Uphoff, O. Fochler, Z. Xu, and C. Greiner, *Phys. Lett. B* **717**, 430 (2012); *Nucl. Phys.* **A932**, 247 (2014).
- [23] R. Sharma, I. Vitev, and B.-W. Zhang, *Phys. Rev. C* **80**, 054902 (2009).
- [24] D. Beavis *et al.*, <https://drupal.star.bnl.gov/STAR/starnotes/public/sn0600>, 2011.
- [25] H. Qiu (STAR Collaboration), *Nucl. Phys.* **A931**, 1141 (2014).
- [26] W. J. Llope *et al.*, *Nucl. Instrum. Methods Phys. Res., Sect. A* **522**, 252 (2004).
- [27] B. Abelev *et al.* (STAR Collaboration), *Phys. Rev. C* **79**, 034909 (2009).
- [28] K. Olive *et al.*, *Chin. Phys. C* **38**, 090001 (2014).
- [29] M. Anderson *et al.*, *Nucl. Instrum. Methods Phys. Res., Sect. A* **499**, 659 (2003).
- [30] W. J. Llope (STAR Collaboration), *Nucl. Instrum. Methods Phys. Res., Sect. A* **661**, S110 (2012).
- [31] A. Hocker *et al.*, *Proc. Sci.*, ACAT2007 (2007) 040.
- [32] N. Borghini, P. M. Dinh, and J.-Y. Ollitrault, *Phys. Rev. C* **63**, 054906 (2001).
- [33] A. Bilandzic, R. Snellings, and S. Voloshin, *Phys. Rev. C* **83**, 044913 (2011).
- [34] H. Masui, A. Schmah, and A. M. Poskanzer, *Nucl. Instrum. Methods Phys. Res., Sect. A* **833**, 181 (2016).
- [35] L. Adamczyk *et al.* (STAR Collaboration), *Phys. Rev. D* **86**, 072013 (2012).
- [36] B. I. Abelev *et al.* (STAR Collaboration), *Phys. Rev. C* **77**, 054901 (2008).
- [37] L. Adamczyk *et al.* (STAR Collaboration), *Phys. Rev. Lett.* **116**, 062301 (2016).
- [38] J. Adams *et al.* (STAR Collaboration), *Phys. Rev. Lett.* **93**, 252301 (2004).
- [39] L. G. Pang, Y. Hatta, X. N. Wang, and B. W. Xiao, *Phys. Rev. D* **91**, 074027 (2015); L. G. Pang, Q. Wang, and X. N. Wang, *Phys. Rev. C* **86**, 024911 (2012); (private communication).
- [40] M. Cacciari, P. Nason, and R. Vogt, *Phys. Rev. Lett.* **95**, 122001 (2005).
- [41] M. Aggarwal *et al.* (STAR Collaboration), *Phys. Rev. Lett.* **105**, 202301 (2010).
- [42] A. Adare *et al.* (PHENIX Collaboration), *Phys. Rev. C* **93**, 034904 (2016).
- [43] R. Esha, M. Nasim, and H. Huang, *J. Phys. G* **44**, 045109 (2017).
- [44] D. Molnar and S. A. Voloshin, *Phys. Rev. Lett.* **91**, 092301 (2003).
- [45] S. Afanasiev *et al.* (PHENIX Collaboration), *Phys. Rev. Lett.* **99**, 052301 (2007).
- [46] D. Banerjee, S. Datta, R. Gavai, and P. Majumdar, *Phys. Rev. D* **85**, 014510 (2012).
- [47] H. T. Ding, F. Karsch, and S. Mukherjee, *Int. J. Mod. Phys. E* **24**, 1530007 (2015).